



SBAS Ephemeris & Ionospheric Corrections and Integrity

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Introduction-SBAS

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Introduction-SBAS



Item	Single Frequency SBAS	DFMC
Service	NPA	CAT-I
Service Signal	GPS L1	GPS L5
Augmentated Signal	GPS L1 C/A	GPS L1 C/A & L5 BDS B1C & B2a
Position Accuracy (95%)	Horizontal: 2.5m Vertical: 4.0m	Horizontal: 1.5m Vertical: 2.0m
Time to Alert	10s	6s
Integrity	2×10 ⁻⁷ /150s	2×10 ⁻⁷ /150s
Continuity	1-8×10 ⁻⁶ /15s	1-8×10 ⁻⁶ /15s
Availability	99%	99.9%



Trimble RTX[™] Satellite Broadcast Frequency Coverage Map

RTXWN (Western North America) 1557.8615 MHz at 600 baud

RTXCN (Central North America) 1557.8150 MHz at 2400 baud RTXEN (Eastern North America) 1557,8590 MHz at 600 baud

RTXSA (Latin America) 1539.8325 MHz at 600 baud



RTXAE (Europe/Africa) 1539.8125 MHz at 2400 baud

RTXAP (Asia/Pacific) 1539.8325 MHz at 600 baud





RTX: Real Time eXtended





RTX : ✓ Position accuracy: H:2.5cm,V:5cm ; ✓ Convergence time : < 20min ; ✓ Service area: Global

RTX FAST :

✓ Position accuracy: H:2.5cm,V:5cm ;

✓Convergence time : < 1min ;
</pre>

✓ Service area: America and Europe







Manufacturer	Coverage	Performance	
Trimble	Global	RTX:H:2.5cm,V:5cm;Convergence Time: < 15min	
	Region	RTX Fast : H:2.5cm,V:5cm ; Convergence Time: < 1min	
OmniSTAR (Trimble)	Global	VBS: Sub-meter ; HP: 5-10cm ; XP: 20-35cm ; G2: Decimeter ;	
TerraStar (Hexagon)	Global	TerraStar-C: 5cm; Convergence Time: 25min ; TerraStar-D: 10cm; Convergence Time: 20min ;	
Veripos (Hexagon)	Global	Apex: H:5cm, V:12cm ; Ultra: H:10cm, V:20cm ;	
StarFIRE	Global	10cm ; Convergence Time: 20min	
QZSS	Global	CLAS: Centimeter; Convergence Time: < 1min; MADOCA: Contimeter: Convergence Time: < 30min;	
A Contraction	Region		
Atlas/ChinaCM	Global	4cm; Convergence Time: 10-40min;	
Hi-RTP	Global	4cm;	





Introduction

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Slow clock correction:

$$\rho_i^j = \rho^0 - Gdeph + cdt_i - cdt^j + \varepsilon$$

$$G = \left[\frac{x^{sat} - x}{\left|\vec{r}^{sat} - \vec{r}\right|}, \frac{y^{sat} - y}{\left|\vec{r}^{sat} - \vec{r}\right|}, \frac{z^{sat} - z}{\left|\vec{r}^{sat} - \vec{r}\right|}\right], \quad deph = \left[dx^{sat}, dy^{sat}, dz^{sat}\right]$$

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- First solve the clock errors of satellites and stations;
- Chose one station as the master station;
- ✓ ignore the orbit error at this step ;
- Satellite clock error shall absorb the orbit radial direction error.
 - This clock results are treat as slow clock corrections





Model for orbit and clock corrections

Slow orbit correction:

$$\hat{X}_{MV} = (\Lambda^{-1} + G^T W^{-1} G)^{-1} G^T W^{-1} z$$

 $\hat{P}_{MV} = (\Lambda^{-1} + G^T W^{-1} G)^{-1}$

- Minimum variance (MV) estimation method is used;
- $\checkmark \Lambda$ Is the prior variance of the ephemeris;
- z is the residual after removing the satellite and station clock errors;
- ✓ ignore the clock error at this step ;
- Carrier phase epoch changes are used to constrain the orbit;
- This orbit results are treat as slow orbit corrections.
- After this, clock errors are solved again as fast corrections.





 $e = (\tilde{P}^{-1} + G^T W G)^{-1} G^T W z$ $P = (\tilde{P}^{-1} + G^T W G)^{-1}$

$$k_{state} = \sqrt{e^T P^{-1} e}$$

- Calculate the residual or remain ephemeris errors e;
- $e^T P^{-1} e$ will follow the chi-square distribution while e is Gaussian type function;
- \checkmark k_{FA} is 4.3 fault alarm rate is 10⁻³
- ✓ The satellite is consider as normal when $k_{state} ≤ k_{FA}$

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UDRE is broadcasted to bound the residual errors after application of ephemeris corrections.

$$P_{broadcast} = \sigma_{UDRE}^2 M_{28} \quad P_{broadcast} = (\frac{1}{5.33})^2 (k_{md} + k_{FA})^2 H_{CR}$$

Consider station i is fault and in the service area:

$$\sigma_{j} = M_{ax}(\sqrt{u_{j}^{T}P_{i}u_{j}}) \quad F_{0} = M_{ax}(\frac{\sigma_{j}}{\sqrt{u_{j}^{T}Pu_{j}}})$$

$$P_{broadcast} = (\frac{1}{5.33})^{2}(k_{md} + k_{FA})^{2}F_{0}^{2}P$$

$$P_{broadcast} = \overline{U}^{T}\overline{U} \quad UDRE = \overline{U}_{4,4}, MT_{28} = \overline{U}/\overline{U}_{4,4}$$

 $\checkmark k_{md}$ is 6.13 when the missed detection rate is 4.5×10^{-10}

 $\checkmark P_i$ is the posterior covariance without station i

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Corrections for precise orbit and clock

- Should be close to zero
- Clock accuracy is about 10cm
- Orbit accuracy is about 40cm
- Fast clock correction is small
- DFMC cancel fast correction







Corrections for broadcast orbit and clock

- Clock accuracy is about 10cm
- Orbit and clock accuracy degrade with satellite growing up
- Fast clock correction is small







- Chi-square test is used to ensure the integrity \checkmark
- Test threshold is 4.3 for 4 parameters \checkmark
- \checkmark UDRE is decreasing while satellite growing up



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Simulation faults of ephemeris

- ✓ First time, add errors on satellite clock(0.5m/s, last 100s,@475200);
- Second time, add errors on satellite orbit(Y, 0.5m/s, last 100s,@475200);



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Position results

- Klobuchar model for ionospheric delay
- ✓ SBAS H<0.8m, V<1.5m
- Improved about 30% and
 20% in Horizontal and
 Vertical.
- With SBAS Iono model will be improved further

1.5 Horizontal Error/m SBAS Error 1 GPS Error 0.5 0 30 5 10 15 20 25 0 2 Vertical Error/m 0 5 20 10 15 25 30 0 Station number

SBAS & GPS position error





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Ionospheric delay is a major error source for SBAS

- Ionospheric delay on each IGP is broadcasted to improve position accuracy;
- While a GIVE is also broadcasted to bound the residual error after ionospheric correction;
- GIVE is inflated by several parameters including $\sigma_{unde\,rsampled}$ to meet the integrity requirements;
- GIVE is a dominant factor of protection levels;

Spatial threat model should be built for GIVE calculation

- Spatial threat model is necessary for GIVE integrity because of the undersampled IPPs during ionospheric irregularity;
 - Spatial threat model is depend on regions, station distribution;
- Data deprivation method is used to generate the threat model;

Planar fit is used to estimate the ionospheric delay at an IGP, assume ionospheric vertical delay can be modeled as a plane.

$$\begin{split} \hat{I}_{v,IGP}(x,y) &= \hat{a}_{0} + \hat{a}_{1}x + \hat{a}_{2}y \qquad I_{v,IPP} = G \cdot x \\ I_{v,IPP} &= \begin{bmatrix} I_{v,IPP} \\ I_{v,IPP} \\ \vdots \\ I_{v,IPP} \end{bmatrix}, G = \begin{bmatrix} 1 & d_{IPP,IGP} \cdot \hat{E} & d_{IPP,IGP} \cdot \hat{N} \\ 1 & d_{IPP,IGP} \cdot \hat{E} & d_{IPP,IGP} \cdot \hat{N} \\ \vdots \\ I & I \end{bmatrix}, x = \begin{bmatrix} \hat{a}_{0} \\ \hat{a}_{1} \\ \hat{a}_{2} \end{bmatrix}, \sigma_{l_{v,IPP}}^{2} = \begin{bmatrix} \sigma_{l_{v,IPP}}^{2} \\ \sigma_{l_{v,IPP}}^{2} \\ \vdots \\ \sigma_{l_{v,IPP}}^{2} \end{bmatrix}$$

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IPP search method for planar fit

- Chose IPP with satellite elevation greater than 5°;
- If IPP number more than N_{target} in the region with radius of R_{min}, then R_{fit} is R_{min};
- Or, expand R_{fit} until it defines a circle that surrounds N_{target} IPPs;
- If R_{fit} attains a maximum value of R_{max} and IPP number still less than N_{target} , but more than N_{min} , a estimation is also performed;



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- Or, the GIVE to be broadcast is set to Not Monitored.

 $R_{\min} = 800 km, N_{t \arg et} = 30, R_{\max} = 2100 km, N_{\min} = 10$



Irregularity detection and GIVE inflation

- Chi-square is used to evaluate the goodness of fit

$$\chi^2 = (I_{v,IPP} - \hat{I}_{v,IPP})^T W (I_{v,IPP} - \hat{I}_{v,IPP}) = I_{v,IPP}^T \cdot W \cdot [I - G^T \cdot (G \cdot W \cdot G^T)^{-1} \cdot G \cdot W] \cdot I_{v,IPP}$$

- With the degree of freedom and false alarm rate a threshold can be determined, if the test value of χ^2 is greater than threshold then the GIVE is set to Not Monitored,
- Or a GIVE based on the formal variance is inflated to guarantee integrity

$$\sigma_{GIVE}^{2} = R_{irreg}^{2} \sigma_{IGP_{k}}^{2} + \max(R_{irreg}^{2} \sigma_{decorr}^{2}, \sigma_{undersampled}^{2} + \sigma_{rate-of-change}^{2} + \sigma_{antenna_bias}^{2}$$
Formal Variance Spatial Threat Temporal Threat Antenna Bias
$$R_{irreg}^{2}(P_{fa}, P_{md}) = \frac{\chi_{1-P_{fa}}^{2}}{\chi_{P_{md}}^{2}}$$
, depend on false alarm rate and miss detection rate

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Undersampled threat

- Ionospheric delay on IGP is estimated with IPPs measured by SBAS monitor stations;
- Local lonospheric irregularities might not be sampled by sparse monitor stations;
- Users might have IPPs within the irregularities but still use surrounding IGPs to estimate the delay which will lead to potential threat of large position error;
- So a spatial threat model should be built and added to GIVE to protect users against such a condition;
- Spatial threat model created based on the historical ionospheric storm data.



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Undersampled threat model examples



- Function of R_{fit} (fit radius) and RCM (Relative Centroid Metric);
- Monotonic function of the two metrics;
- Created with data deprivation method.

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Data deprivation method

- Threat region: the region between the centers of the surrounding ionospheric grid squares for any IGP;
- Annular deprivation:
- separate out data in successive annuli, IPPs on an annulus (red plots) are not used for fit;
- The width of each annulus is 200km;
- The inner radius of annuli changes from 200km (exclude the threat region) to 2000km;
- IPPs in threat region are as virtual users.









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Data deprivation method

- Three quadrant deprivation:
- IPPs in three quadrants are excluded form fit algorithm;
- IPPs in the remain quadrant are used to perform the planar fit;
- the cutoffs are done at every 100km within a 500km range in four directions.
- Residuals between measurements and estimation from planar fit of IPPs in the threat region are prepared for creating the threat model.



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Calculation of $\sigma_{undersampled}^{raw}$

 $I_{\scriptscriptstyle v, \rm IPP}$:is ionospheric delay measurement of an IPP

- K : is nominal 5.33
- R_{irreg} : is a inflation factor from chi-square test
- $\hat{I}_{v,IPP}$: is planar fit estimation

 $\sigma_{\rm GIVE}$: is formal variance form planar fit of the IGP

R_{fit} and RCM metric

$$\begin{bmatrix} 1 \\ d_{cent,x} \\ d_{cent,y} \end{bmatrix} = \frac{G^T \cdot W \cdot 1}{1^T \cdot W \cdot 1} \qquad RCM = \sqrt{d_{cent,x}^2 + d_{cent,y}^2} / R_{fit}$$





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26 stations in Crustal Movement Observation Network of China are selected.





Data chosen according to Kp and Dst index

Four days of data in 2015 will be used in the analysis.



Day	Kp index	Dst index/nT
76	7.7	-223
174	7.7	-204
280	6.0	-124
354	6.3	-155



Day of 2015

Data source is from: https://omniweb.gsfc.nasa.gov/form/dx1.html



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Annular Deprivation VS Three Quadrant Deprivation (Day 176)



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With VS without Irregularity detection





With detection



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Threat model of different latitude









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System development



GPS PRN:147&148

System provides PPP and SBAS (Non SOL) service.







System development

WAN

147:5:

⊗





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Thank you !